A brief recall

aerosol (satellite) retrievals

 Thomas, Kolmonen, Lang, Sayer, Xue, Wagner, Povey, Duan, Stap, deVries, Yoon, Fan, Mazolla, Sogocheva, Sano, Penning de Vries

- new ideas
- evaluations
- uncertainties
- by association

new retrieval concepts



Use of corrected reflectance

AOD evaluation

here for ...

AATSR-ADV Level 3 (1° x 1° grid) for 2008 against AERONET (using several AeroCom tools)







estimating pixel uncertainty



general typing by 'association'



Aerosol types: BB – biomass burning, DD – desert dust, BIO – (sec.) biogenic, URB – (sec.) urban, VOG – volcanic sulfate, AGED – aged or transported, SS – sea salt (na – not assessed)

direct aerosol (rad. transfer) impacts

- Tomasi, Arola, Doppler, Kinne
 - global
 - regional
 - local
 - instantaneous

global distributions



MODIS +CERES → clear-sky forcing





forcing sensitivities





cloud (satellite retrievals)

,CO,

- Hollmann
- Devasthale
- Thomas
- Grosvenor

- MODIS (Aqua+Terra) [DWD]
- AVHRR (NOAA-15, 16, 17) [DWD]

- AATSR [RAL] **cloud cci**

cloud top pressure



MODIS mean droplet concentration for liquid water clouds >80% 1x1 cloud fractions



aerosol-clouds

- Modeling
 - Quaas
 - Rosenfeld
 - Chang
 - Neubauer
- Observations
 - Devasthale
 - Lelli
 - Costantino

AVHRR / MODIS GOME / SCIAMACHY POLDER/ MODIS / CALIPSO

Closed cells

Open cells









Can aerosols close open cells?



Rosenfeld et al., 2012

inferred ... -1.4 W/m2 (+/- 0.3) cooling



-1.2W/m2 (+/- 0.4)

- Uncertainty sources of anthropogenic aerosol effect
- Natural/anthropogenic emissions most important
- Aerosol processes cause 14% of variance
- Low-level stratified clouds from ISCCP data
- anthropogenic aerosol indirect effect (top) and uncertainty (bottom)



pollution ~ cloud-top height



cloud-top height trends (O2-band)

Significant trend ß [m/yr]



pollution ~ 1/precipitation MODIS/CALIPSO



looking ahead

- the power of polarization
 - Breon
 - Hasekamp
- new retrievals
 - Litvinov
 - Di Nioa
- new sensors / opportunities
 - Sano
 - Davis
 - Marbach
 - Sayer
 - Hasekamp

19:23UTC

Step and stare views

29° forward view

discontinuity in fringe positions indicates change in droplet size

29° backward view

Intensity (445, 555, 660)

19:25UTC

smaller drops

larger drops

Intensity (445, 555, 660)

DOLP (470, 660, 865)

Drimary bow

DOLP (470, 660, 865)

GRASP approach (Dubovik et al, 2011) simultaneous inversion of a large group of pixels within one or several images

spatially smooth, spectrally dependent AOD

size distribution (shapeindependent):

- d*V*/dln*r* - volume size distribution in total atmospheric column;

- size distribution is modeled using 22 size bins (0.05 $\leq r \leq$ 15 μ m);

- size distribution is **smooth**

AEROSOL shape and composition (in the total atmospheric column):

- <u>randomly oriented homogeneous</u> <u>spheroids;</u>
- aspect ratio distribution N(ε) is fixed to that retrieved by Dubovik et al. 2006
- $1.33 \le n \le 1.6$; $0.0005 \le k \le 0.5$
- *n* and *k* smooth, spectrally
 <u>dependent</u>

3MI Channels

Channel centre and width	Polarisation	Optical head
410 nm 20 nm	Yes	
443 nm 20 nm	Yes	
490 nm 20 nm	Yes	
555 nm 20 nm	Yes	VNIR Optical head
670 nm 20 nm	Yes	
763 nm 10 nm	Νο	
754 nm 20 nm	No	
865 nm 40 nm	Yes	
910 nm VNIR 20 nm	No	
910 nm SWIR 20 nm	No	
1370 nm 40 nm	Yes	SWIR Optical
1650 nm 40 nm	Yes	head
2130 nm 40 nm	Yes	



Slide: 25



SGLI VIS and near-IR







October 2011 Post Acoustics Alignment

VNR non Polarized Obs. (NP)

- 3 telescopes with 24deg FOV realize the total 70 deg FOV Observation (1,150km)
- Wide wavelength range Observation from **380** to 869 nm.

VNR Polarized Obs. (PL)

- 2 telescopes with 55 deg FOV each for 674 and 869 nm Observation.
- AT tilting mechanism for + / -45deg
- 55 dea FOV with 45 dea tilting



Courtesy of Dr. Tanaka (JAXA/EORC)

The (near) future: VIRS



First VIIRS global image: 24th November 2011, courtesy of NASA NPP team

- Visible Infrared Imaging Radiometer Suite (VIIRS) launched on Suomi-NPP in late 2011
- Similar to MODIS (for aerosol purposes), but:
 - 3,000 km swath width (no gap between orbits)
 - 'Bowtie effect' (pixel size increase across swath) much smaller than in MODIS
 - 750 m pixel size
- Current available products are distributed by NOAA, for operational purposes
 - NASA has recently put out a call for proposals to 'continue the EOS heritage'

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Flight	Units 1 and 2 Ins	trument Specifications	
Orbit:		833 km polar sun-synchronous	
Swath:		>3,000 km (±56 degrees about nadir)	
Scanning:	Rotating telesco	ating telescope with dual-sided, half-angle mirror	
Size:		135 x 148 x 89 cm ³	
Spectral Coverage:		0.4 to 12.5 μm	
Number of Bands: Visible/Near Infrared Mid-wave Infrared Long-wave Infrare	ed: I: d:	9, plus day/night band 8 4	
Resolution: Radiometric (16 ba Imaging (5 bands): Day/Night Band:	ands): :	0.742 km nadir, 1.6 km EOS 0.371 km nadir, 0.8 km EOS 0.742 km constant across scan	
Mass:		270 kg	
Power:		170 W	
Data Rate:		8 Mbps (avg.) / 10.5 Mbps (max.)	



Figure 1. High level VIIRS Flight Unit 1 and Flight Unit 2 instrument characteristics with photo of FU1 being integrated onto the NPP spacecraft at Ball Aerospace. Photo courtesy Ball Aerospace.

SPEX Spectro-polarimeter

- Innovative measurement concept: spectral modulation
- Linear polarization parameters encoded in radiance spectrum by passive optical



Air **MSP**





Spectral bands:

355, 380, 445, 470*,555, 660*, 865*, 935 nm (*polarimetric)



The AirMSPI camera flies in the nose of NASA's ER-2 aircraft (20 km flight altitude)

AirMSPI is mounted in a gimbal for multi-angle viewing between ±67°

what I liked

- actual discussions
 - aerosol absorption (there ARE ideas to follow)
 - (level2) pixel accuracy is becoming a standard!
- overall set-up
 - long breaks
 - time to talk, to exchange ideas

finally

- **BIG** thanks to the organizers
 - Alexander
 - Johannes
 - Gerrit